

HISTORY OF SETI AND THE SEARCH FOR HABITABLE WORLDS



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class 2

SETI

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graph TD; SETI[SETI] --> Technosignatures[Search for Technosignatures]; SETI --> Exoplanet[Exoplanet searches]; Technosignatures --> Radio[Radio SETI]; Technosignatures --> Optical[Optical SETI]; Exoplanet --> Biosignatures[Search for biosignatures]; Biosignatures --> Astrobiology[Astrobiology];
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Search for Technosignatures

Exoplanet searches

Radio SETI

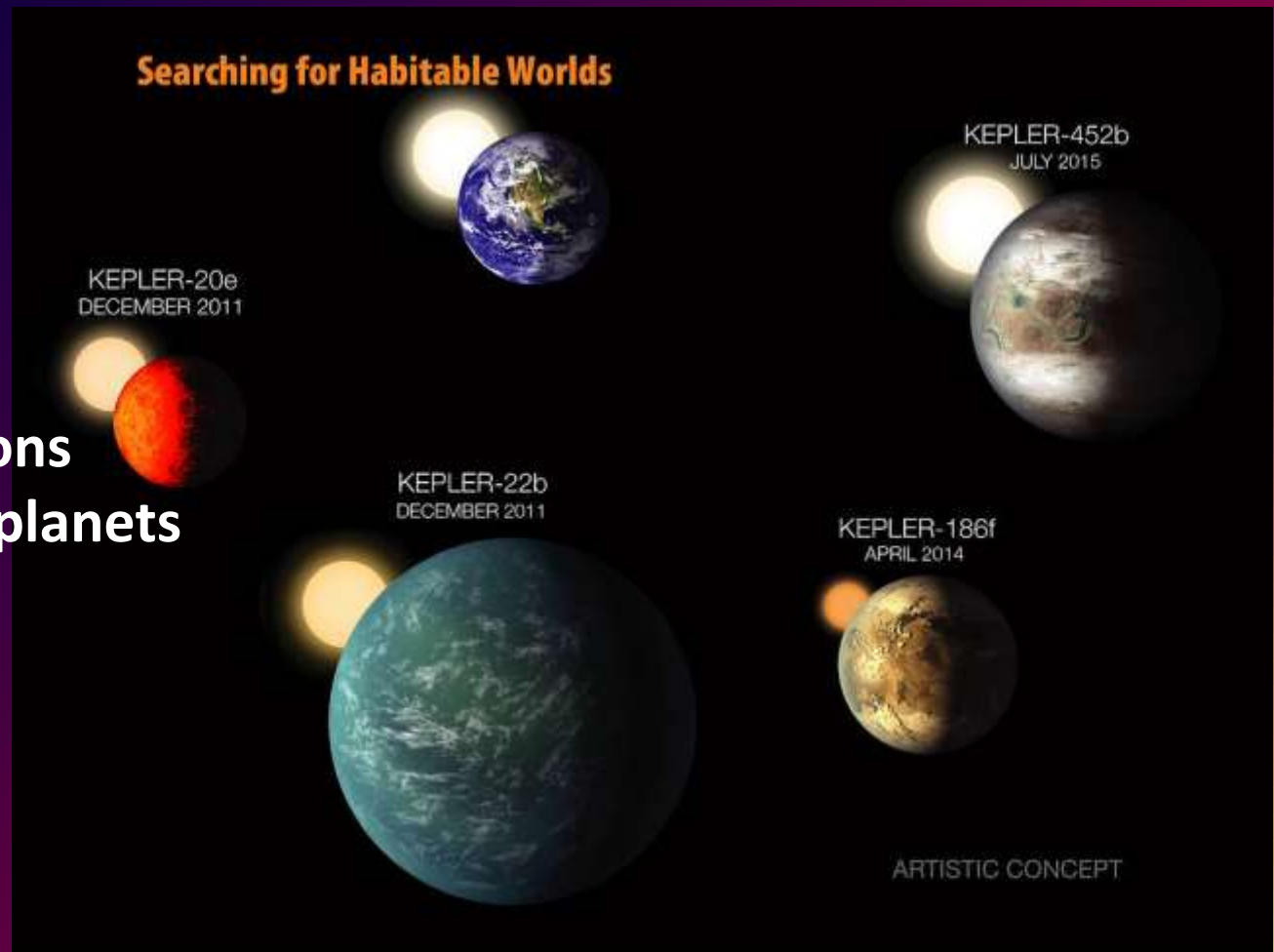
Optical SETI

Search for biosignatures

Astrobiology

Picking up where we left off last week...

One aspect of the search for Ets is searching for habitable worlds around other stars, habitable being defined as able to support liquid water on their surface.



Kepler and K2 Missions
found thousands of planets
orbiting other stars
in our galactic
neighborhood.

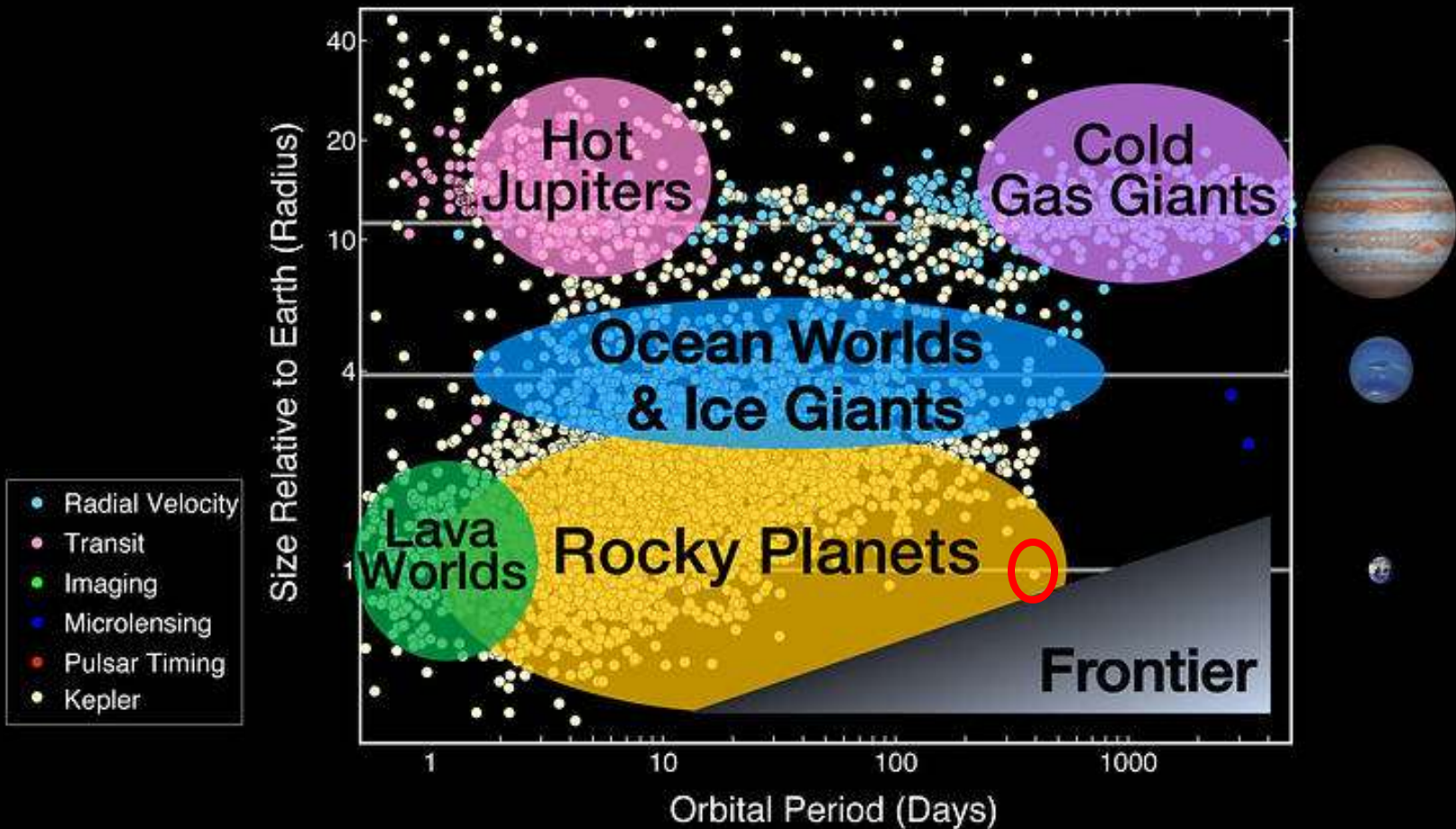
Region of our galaxy imaged
by Kepler and K2 Missions,
looking for exoplanets



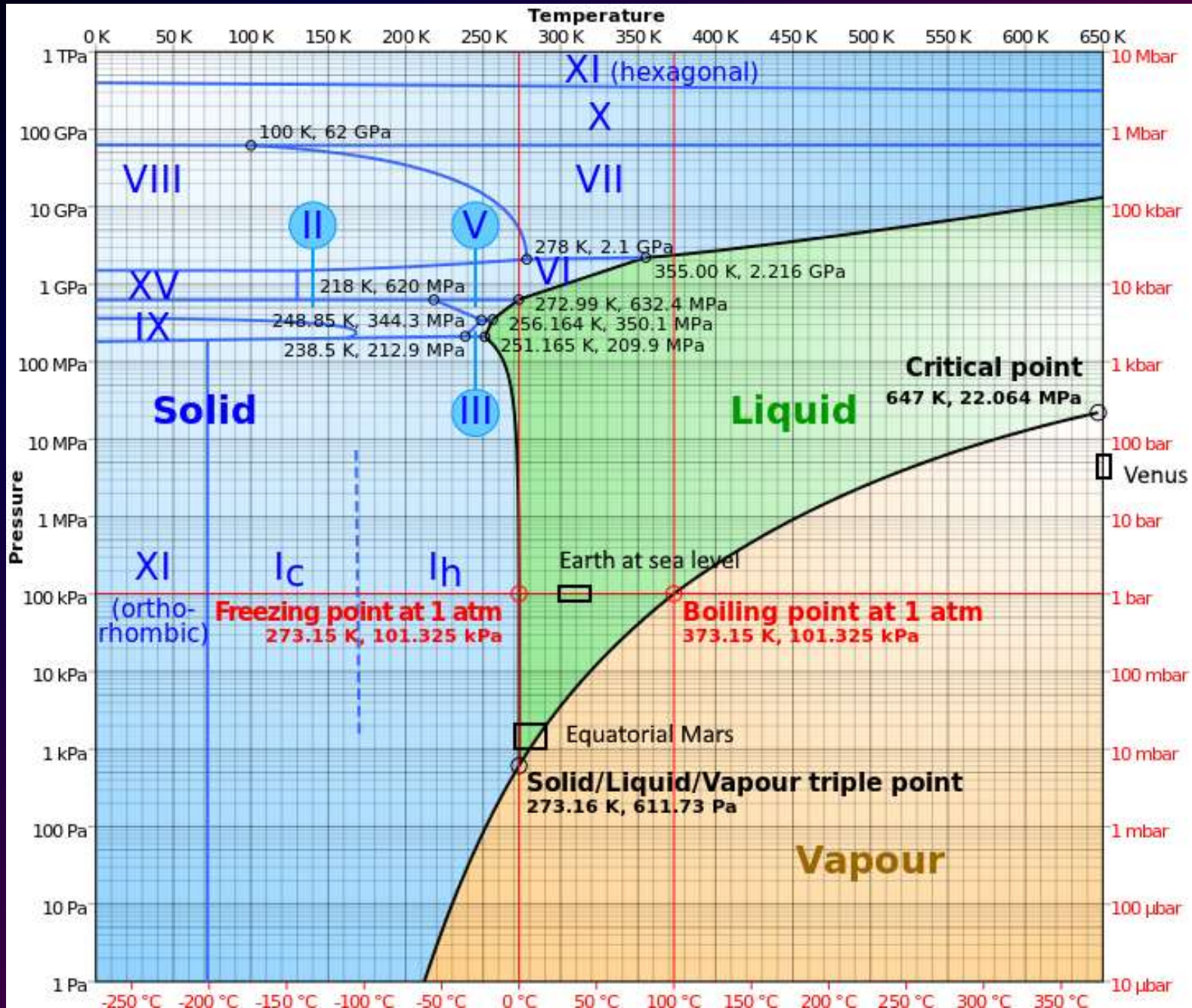
[https://www.youtube.com/watch?
time_continue=2&v=3yij1rJOefM](https://www.youtube.com/watch?time_continue=2&v=3yij1rJOefM)

Types of planets discovered by Kepler and K2 compared to Earth, Neptune, and Jupiter

Exoplanet Populations



PHASE DIAGRAM FOR WATER



MICHAEL H. HART

ICARUS 37, 351-357 (1979)

Our Sun →

TABLE IV
HABITABLE ZONES ABOUT MAIN-SEQUENCE STARS^a

Stellar mass (M/M_{\odot})	Approximate spectral type	Continuously habitable zone		
		r_{inner} (AU)	r_{outer} (AU)	width (AU)
1.20	F7	1.543	1.630	0.087
1.15	F8	1.370	1.454	0.084
1.10	F9	1.221	1.292	0.071
1.05	G0	1.083	1.143	0.060
1.00	G2	0.958	1.004	0.046
0.95	G5	0.840	0.874	0.034
0.90	G8	0.732	0.755	0.023
0.85	K0	0.634	0.649	0.015
0.80	K2	0.542	0.551	0.009
0.75	K4	0.460	0.463	0.003
0.715	K5	0.407	0.407	—

HABITABLE ZONES defined for main sequence stars:

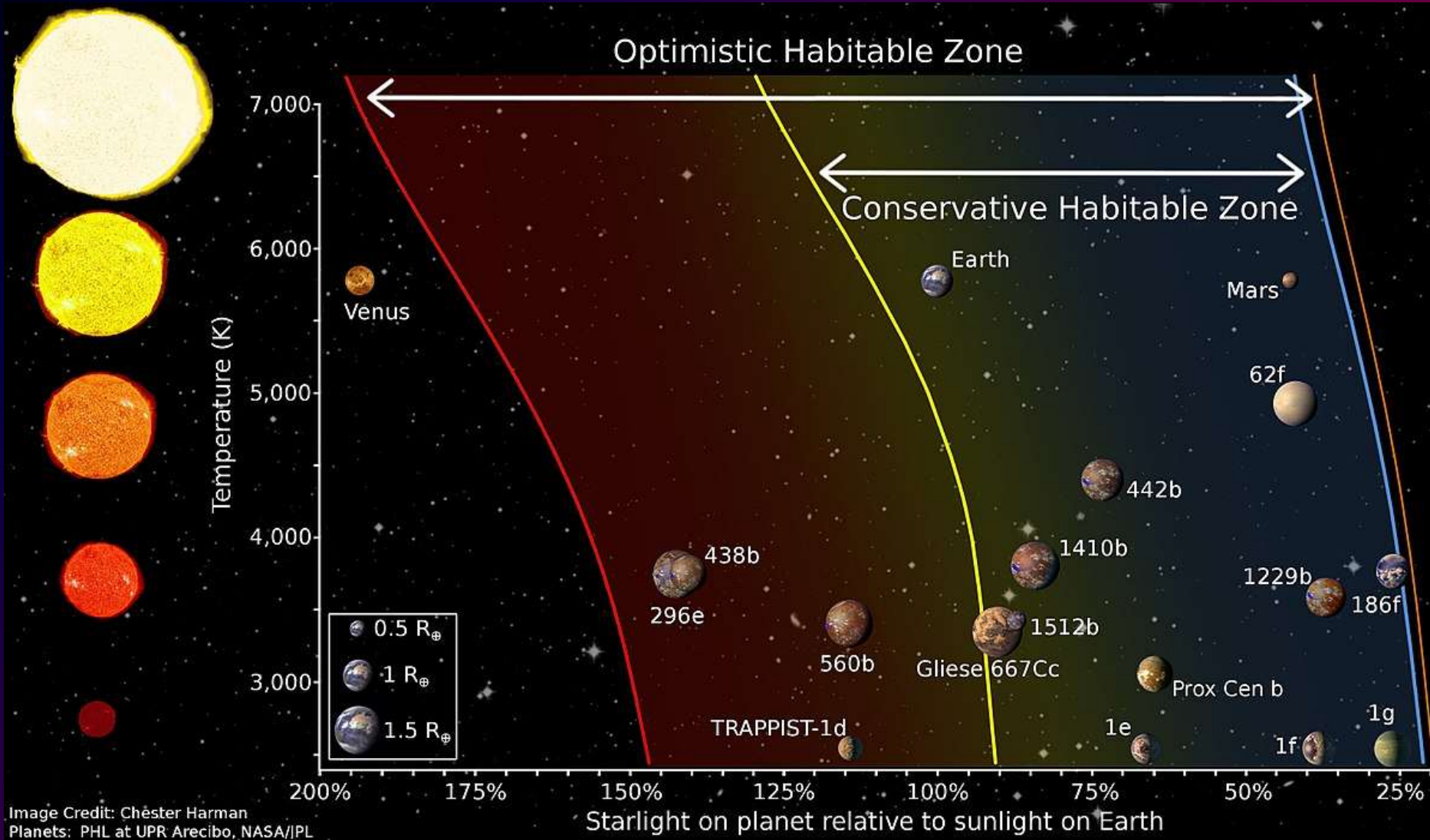
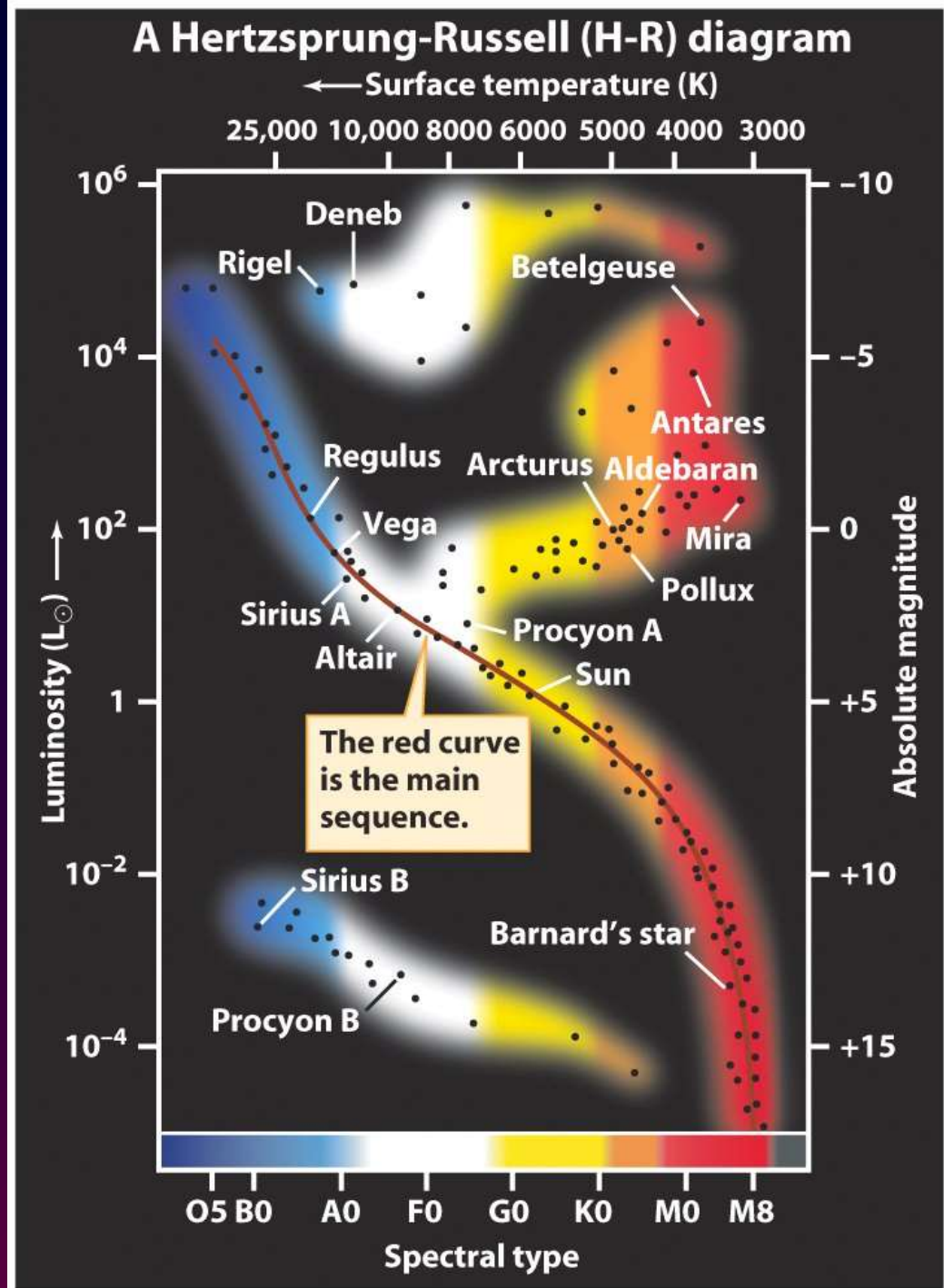


Image Credit: Chester Harman
Planets: PHL at UPR Arcibo, NASA/IPL

Main Sequence:
Stars that are
fusing hydrogen
into helium in
their cores

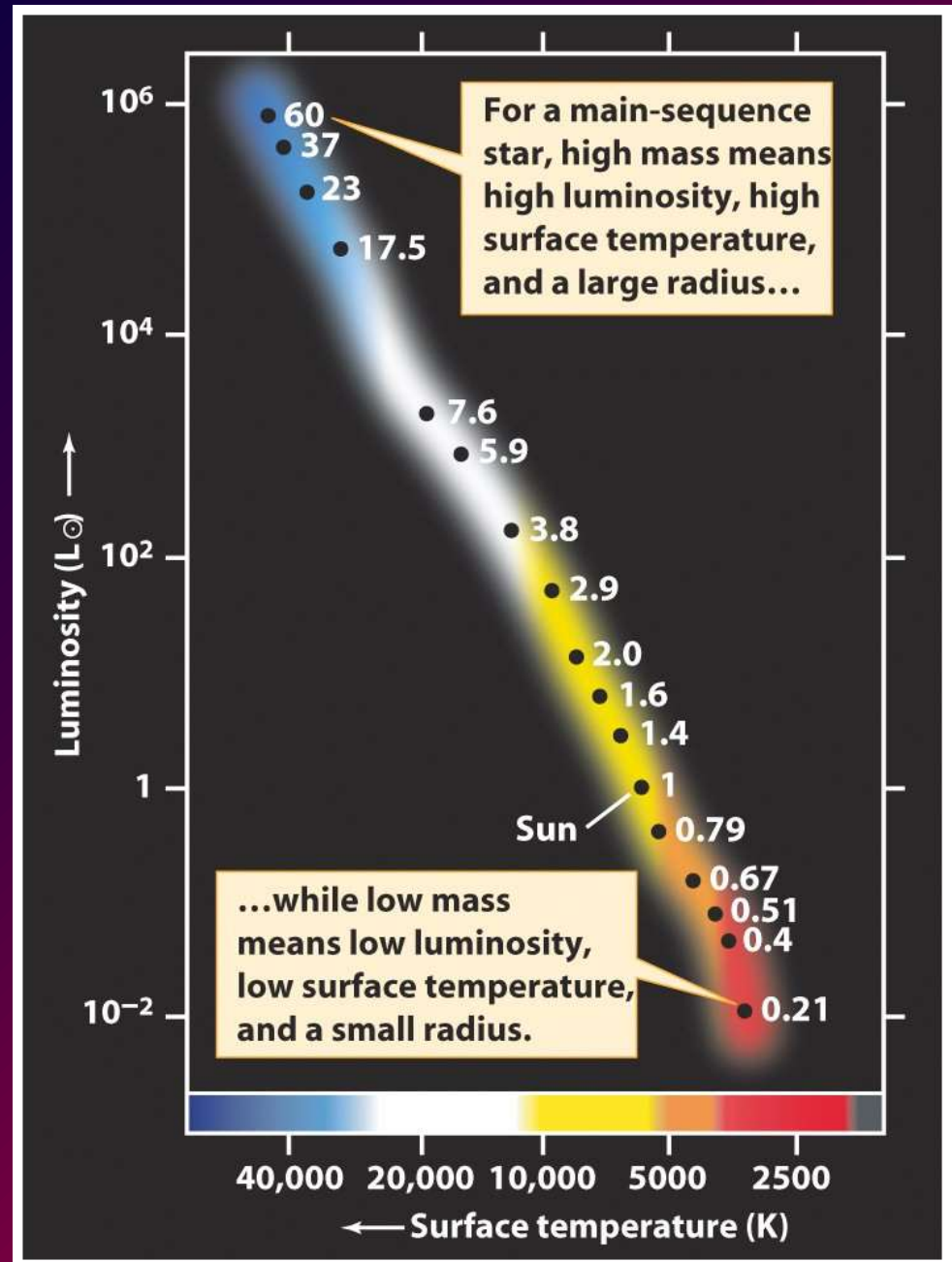


Greater mass means greater central pressure & temperature and greater luminosity...



Greater core pressure increases the RATE of nuclear reactions, resulting in greater luminosity.

Luminosity



Approximate relationship between

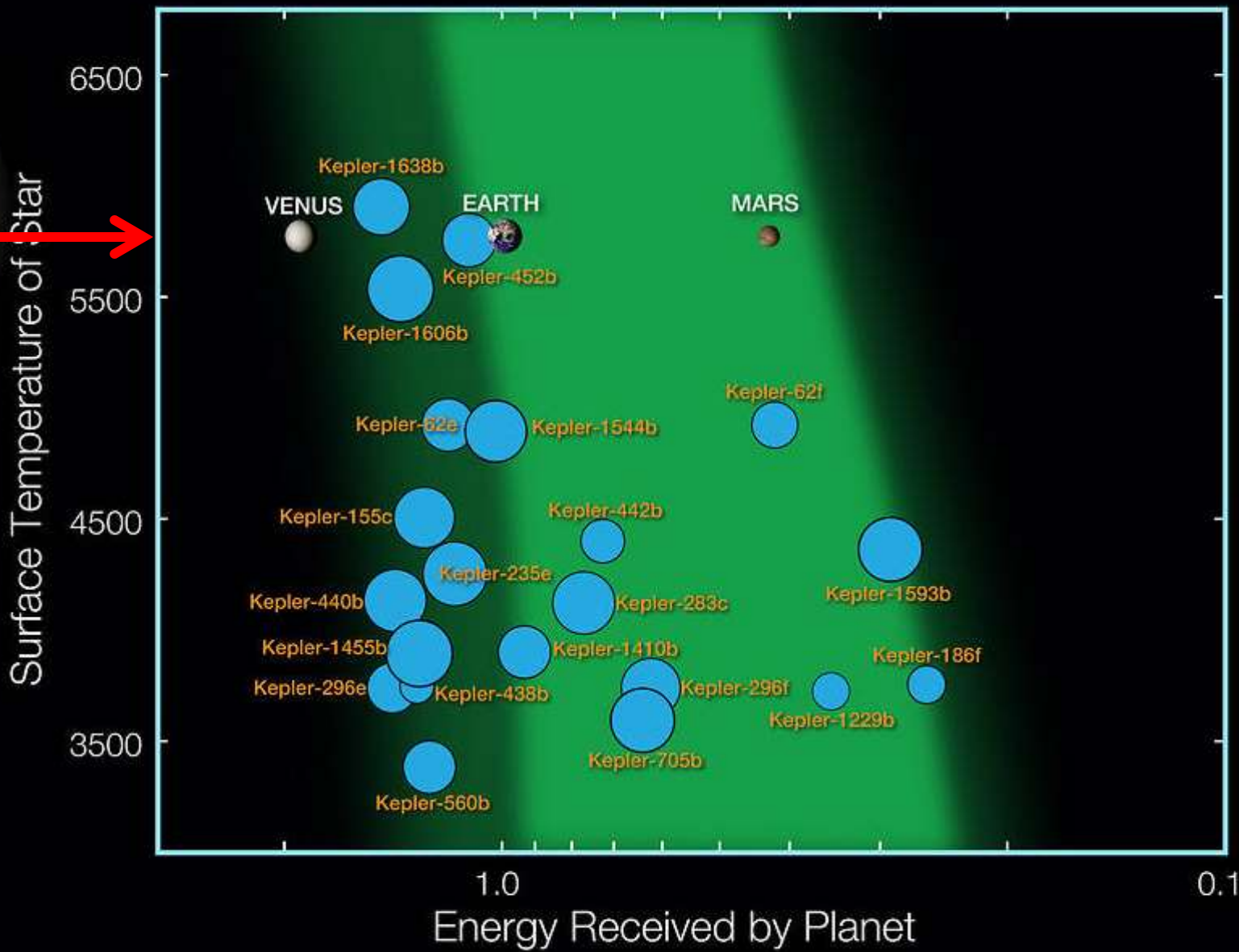
Mass and Luminosity:

$$L \approx M^{3.5}$$

for main sequence stars

Some of the Kepler planets in the habitable zones of their stars

Sun

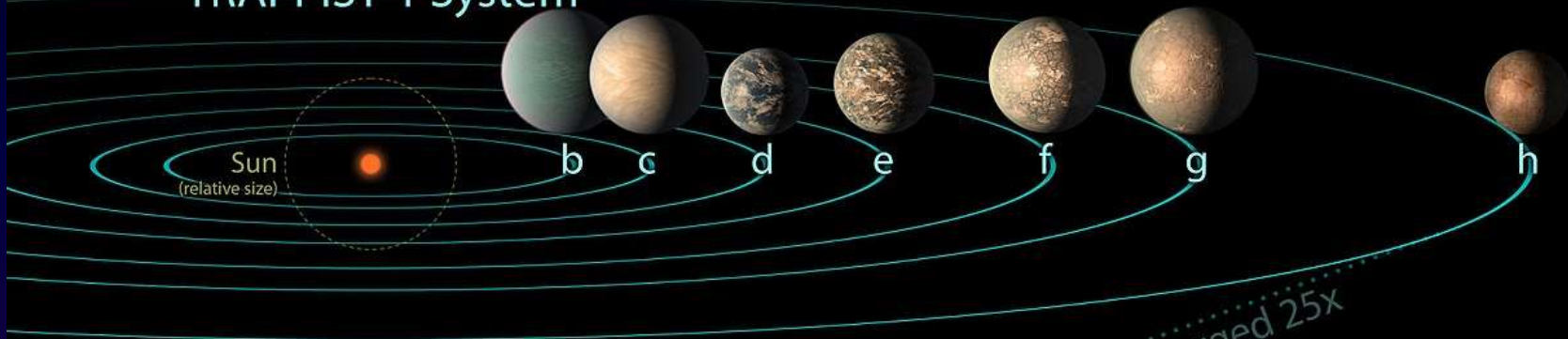


In 2015 the Transiting Planets and Planetesimals Small Telescope (TRAPPIST) discovered a 7-planet system of rocky exoplanets around a small red dwarf, ~ 40 ly away.

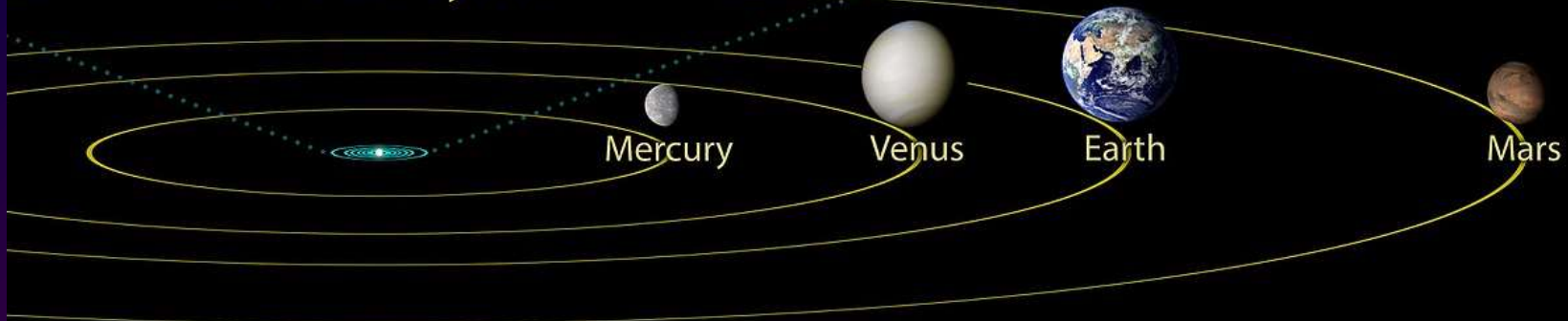
Jupiter & Major Moons



TRAPPIST-1 System

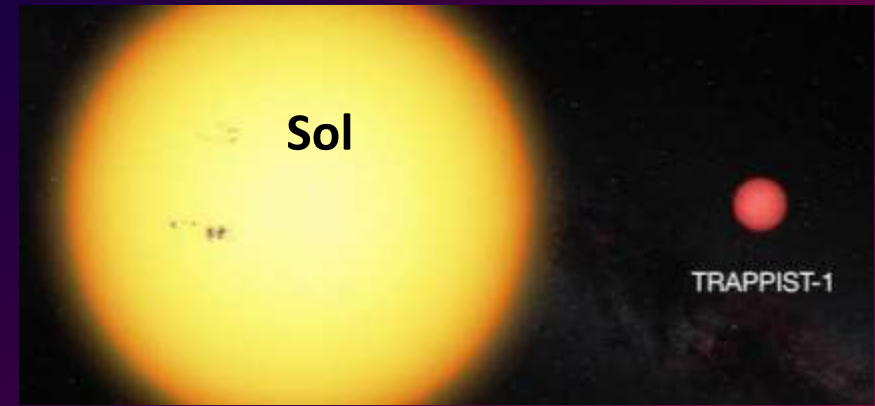


Inner Solar System

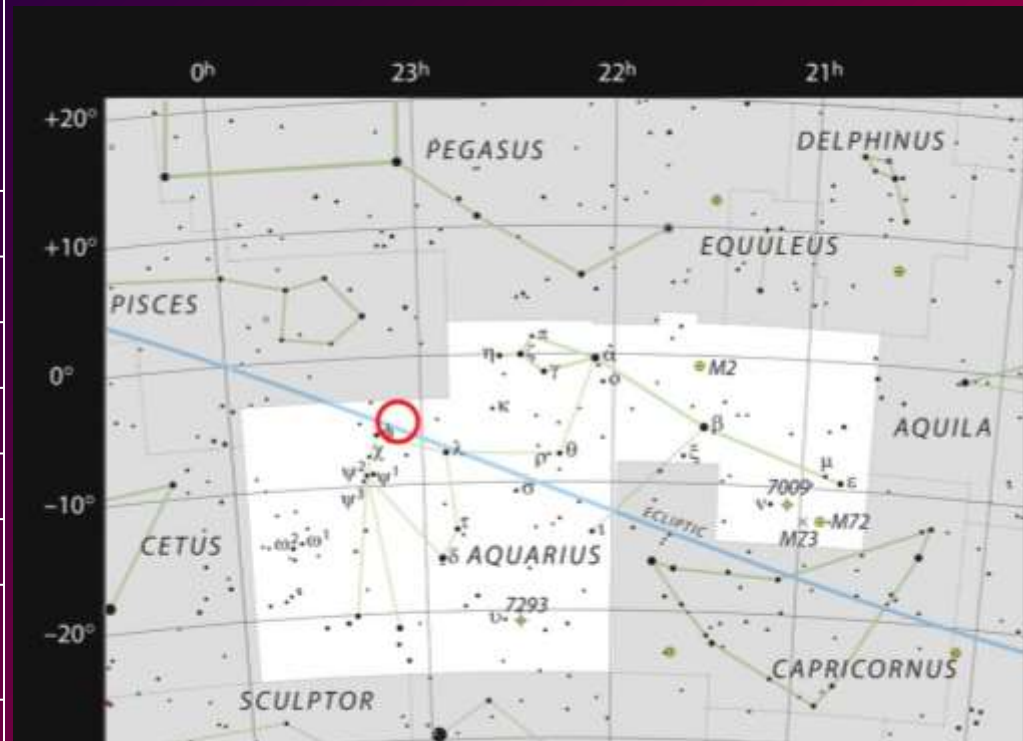


<https://www.nasa.gov/feature/jpl/new-clues-to-trappist-1-planet-compositions-atmospheres>

comparison with our Sun

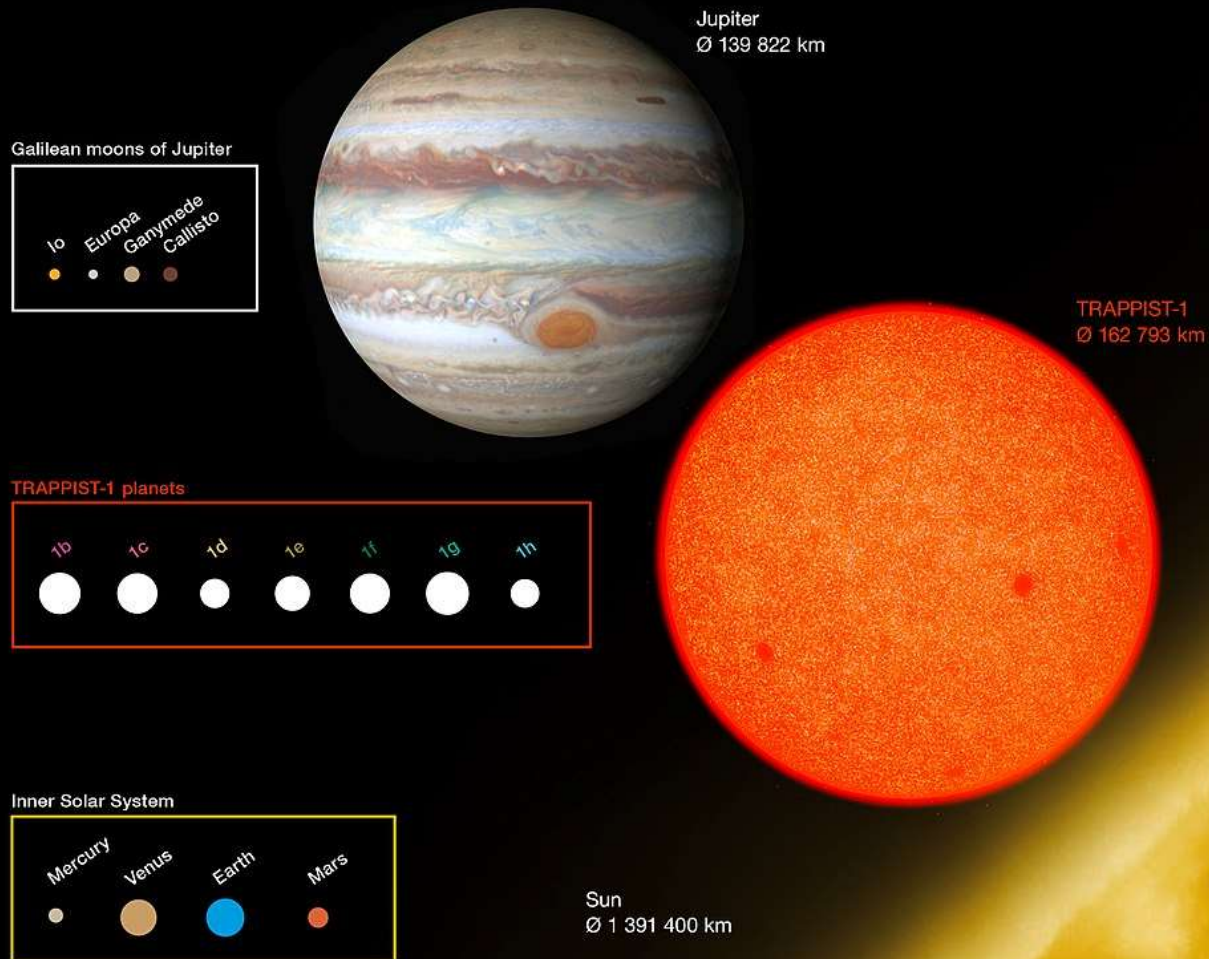


Name	2MASS J23062928-0502285
Alternate name	TRAPPIST-1
Right ascension	$\alpha = 23^{\text{h}} 06^{\text{m}} 29.28^{\text{s}}$
Declination	$\delta = -05^{\circ} 02' 28.5''$
Constellation	Aquarius
Apparent magnitudes	$V = 18.80 \pm 0.08$, $R = 16.47 \pm 0.07$, $I = 14.0 \pm 0.1$, $J = 11.35 \pm 0.02$, $K = 10.30 \pm 0.02$
Parallax	82.4 ± 0.8 mas
Distance	12.14 ± 0.12 pc
Mass	$0.089 \pm 0.006 M_{\odot}$
Radius	$0.121 \pm 0.003 R_{\odot}$
Density	$51.1_{-2.4}^{+1.2} \rho_{\odot}$
Effective temperature	2511 ± 37 K
Luminosity	$0.000522 \pm 0.000019 L_{\odot}$
Metallicity [Fe/H]	$+0.04 \pm 0.08$
Age	7.6 ± 2.2 Gyr



Size Comparison

between TRAPPIST-1 system, Galilean moons of Jupiter and the inner Solar System



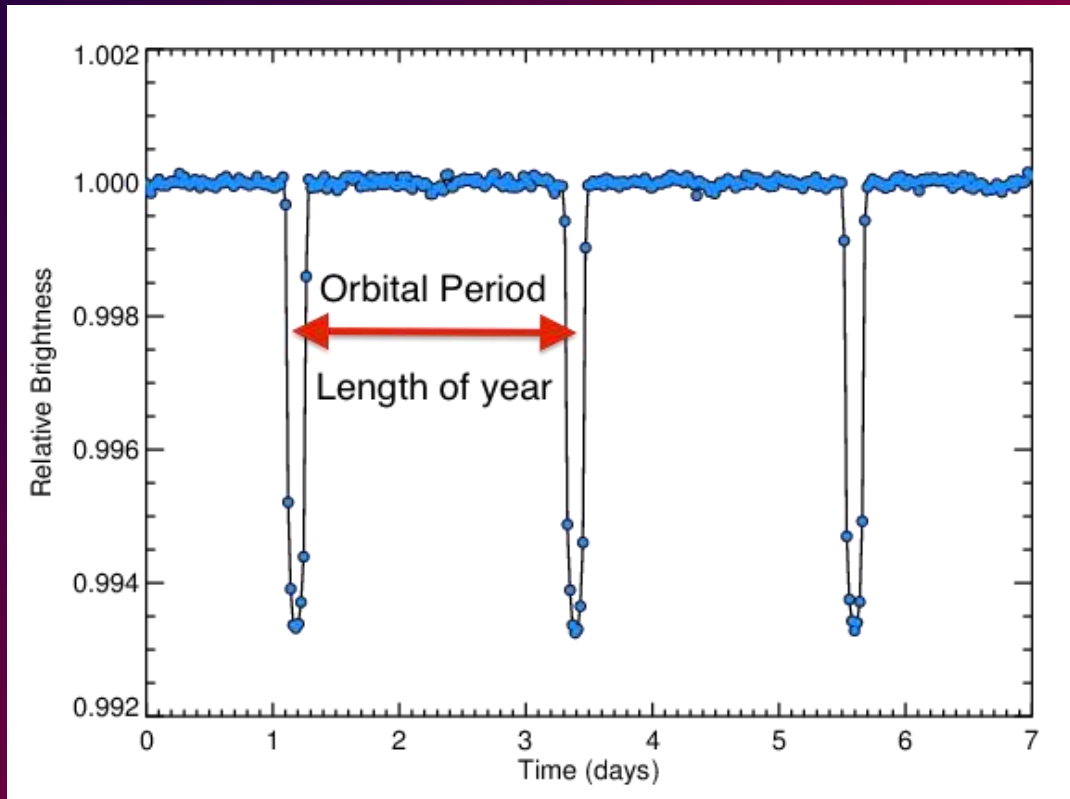
SPECULOOS Southern Observatory

The Search for habitable Planets EClipsing ULtra-cOOl Stars — Finding Earth-like planets around tiny, dim stars using the transit method



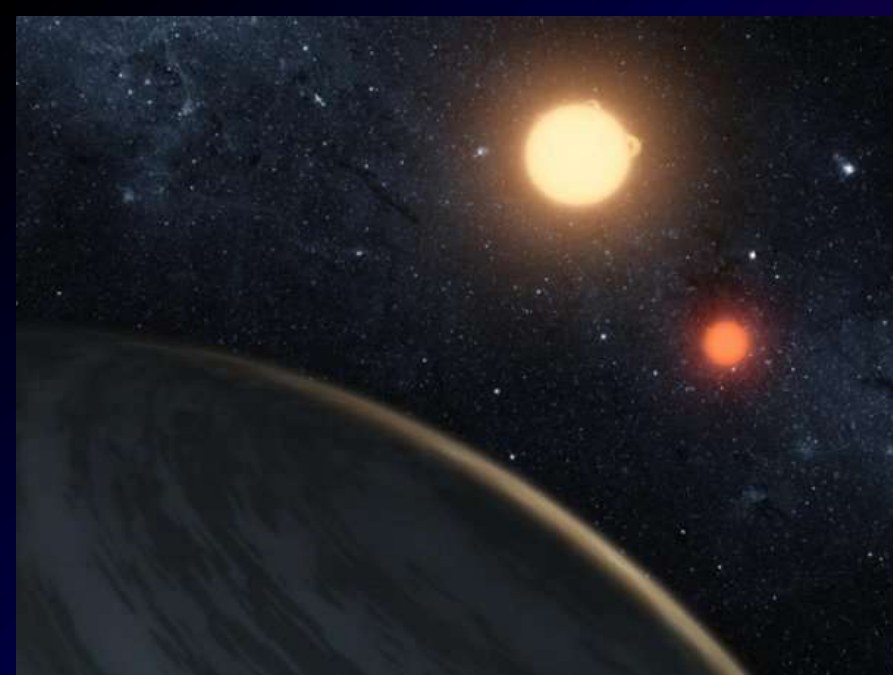
How to find exoplanets:

One method is by observing the light curve of a star. There is a decrease in the brightness observed when a planet transits in front of the star, from our viewpoint.



Kepler 1647b – a real circumbinary planet (artist's rendition)

Detected by the transit method



fictional Tatooine



3,885 confirmed exoplanets
as of Jan. 17, 2019

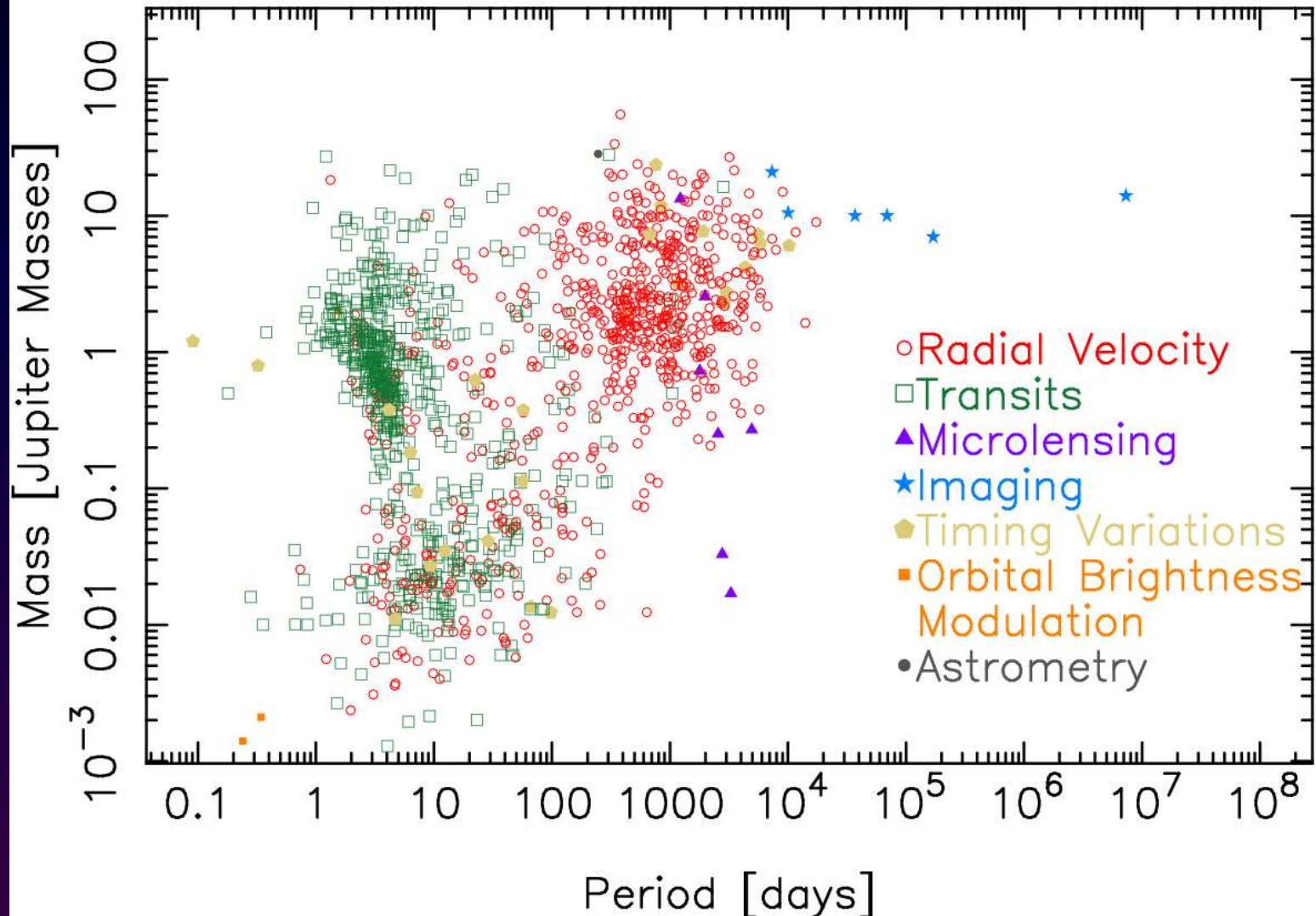
NASA Exoplanet Archive

<https://exoplanetarchive.ipac.caltech.edu/>

Mass – Period Distribution

17 Jan 2019

exoplanetarchive.ipac.caltech.edu



SEARCHING FOR ALIEN LIFE

Astrobiologists are fine-tuning the list of substances that, if spotted on a planet orbiting another star, could constitute evidence of extraterrestrial life.

LIFE AS WE KNOW IT

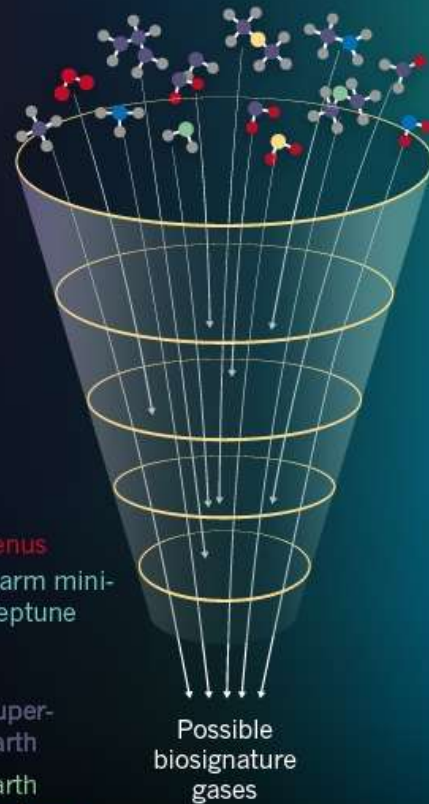
One method is to study a star's light for the chemical imprint of gases that may have been formed by living organisms.



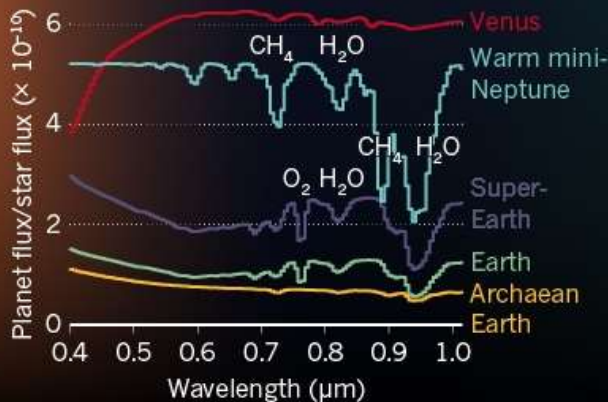
LIFE AS WE DON'T

Another approach is to evaluate a huge range of molecules, winnowing them down on the basis of factors such as stability and detectability.

All small molecules

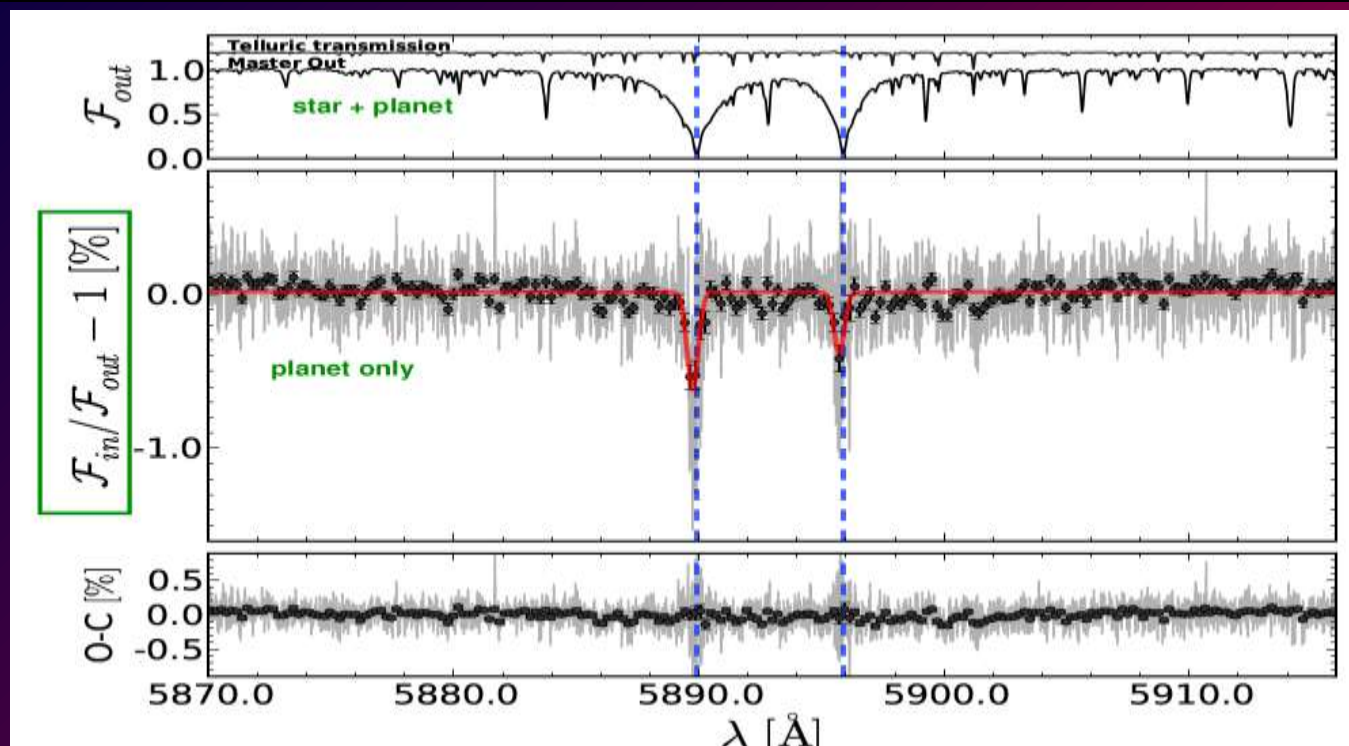
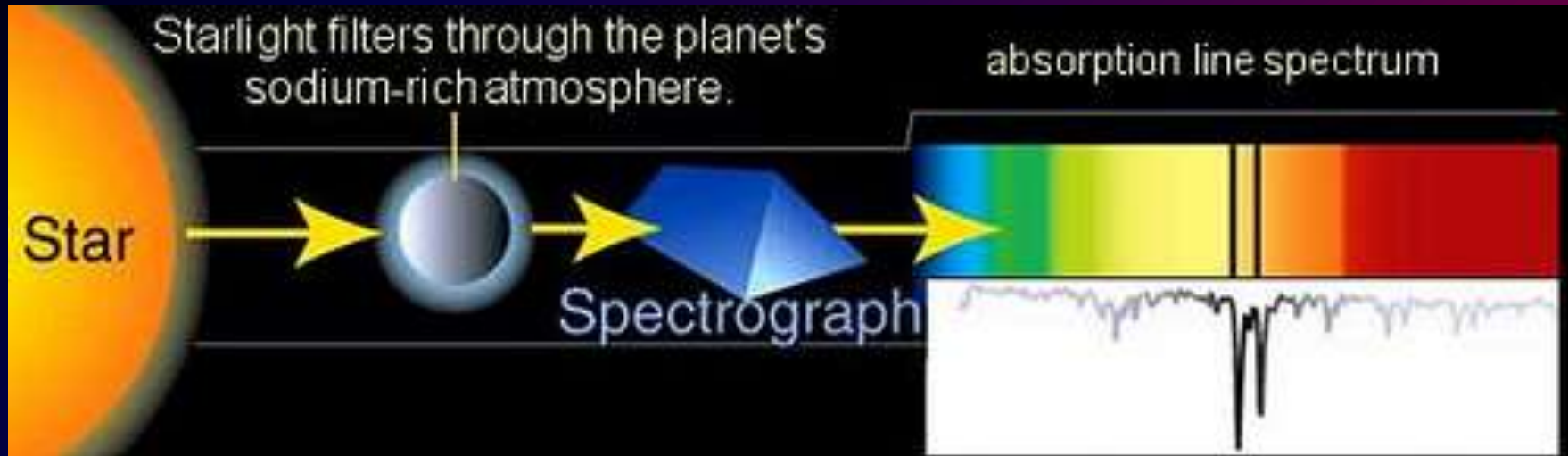


Changes in the starlight transmitted through the planet's atmosphere reveal gases within.



Continuing search for exoplanets extends to the search for biosignatures in the atmospheres of exoplanets

Detecting elements and compounds in the atmospheres of exoplanets by using absorption spectra. Example:



A BRIEF HISTORY OF SETI



Aliens made in our image...

At this very minute, with almost absolute certainty, radio waves sent forth by other intelligent civilizations are falling on the earth. A telescope can be built that, pointed in the right place, and tuned to the right frequency, could discover these waves. Someday, from somewhere out among the stars, will come the answers to many of the oldest, most important, and most exciting questions mankind has asked.

*–Frank D. Drake,
(Intelligent Life in Space
The MacMillan Co.)*

First organized SETI: Project Ozma

1960

Cornell University, NY
Astronomer Frank Drake



National Radio Astronomy Observatory
in Green Bank, West Virginia

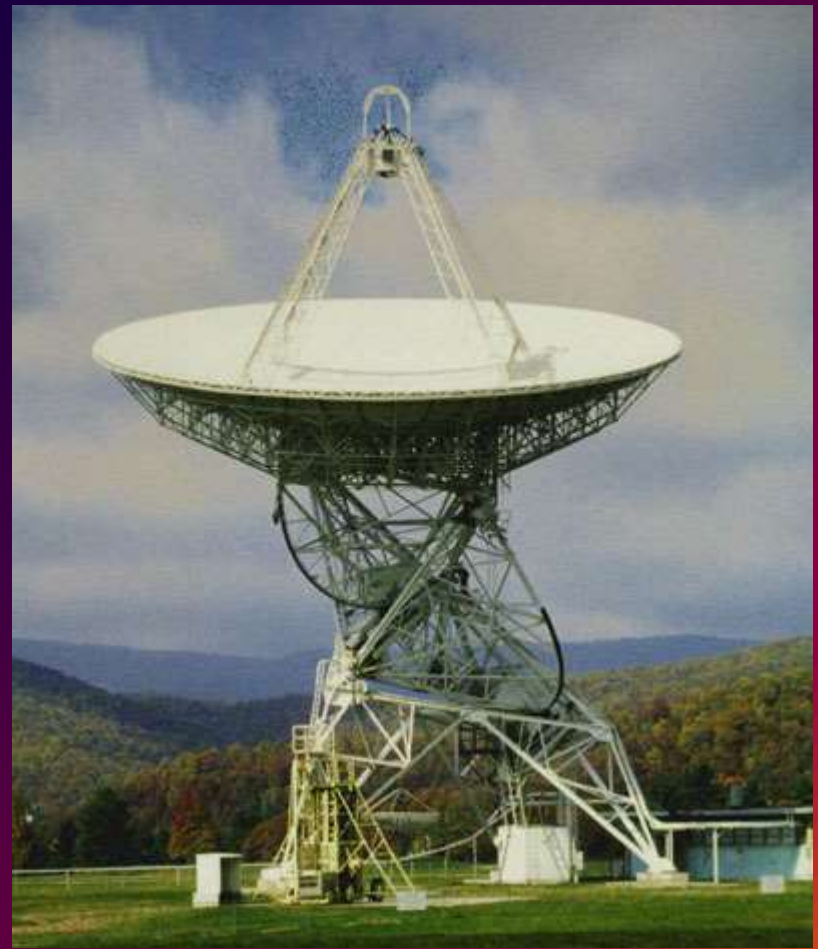
Objective: Search for signs of intelligent
life in distant planetary systems

Method: Scan a 400 kHz bandwidth for
21 cm (1420 MHz) radio frequency emission

21 cm is the frequency of spontaneous spin
flip of a hydrogen atom.

It was thought that this would be familiar to any civilization trying to transmit.

4 months of taking data – no signal found



PROJECT CYCLOPS

A Design Study of a System for Detecting Extraterrestrial Intelligent Life

(NASA-CR-114445) PROJECT CYCLOPS: A
DESIGN STUDY OF A SYSTEM FOR DETECTING
EXTRATERRESTRIAL INTELLIGENT LIFE
(Stanford Univ.) 253 p

N73-18822

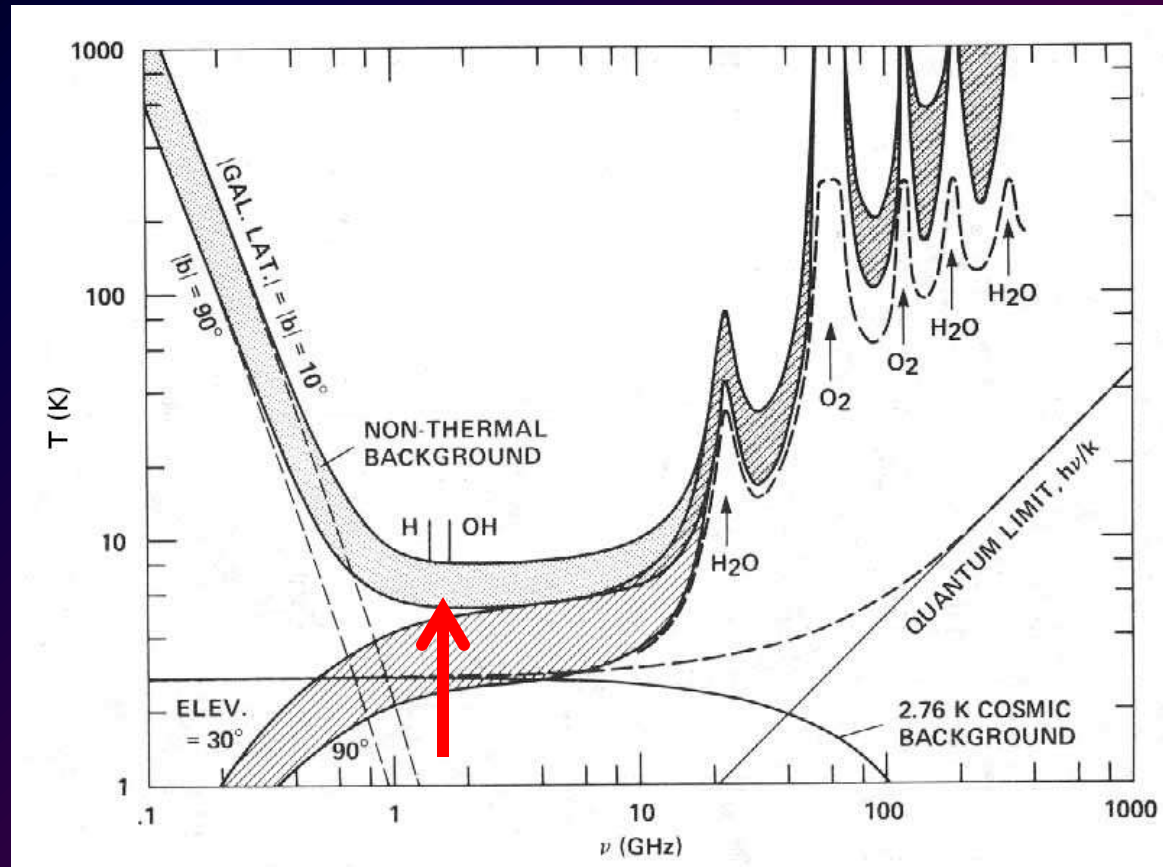
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Unclas
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1971 NASA project
to design an optimal
SETI

using optimal
frequency window
in the microwave
region of the EM
spectrum



“Nature has provided us with a rather narrow quiet band in this best part of the spectrum that seems especially marked for interstellar contact. It lies between the spectral lines of hydrogen (1420 MHz) and the hydroxyl radical (1662 MHz). ...these two emissions of the disassociation products of water beckon all water-based life to search for its kind at the age-old meeting place of all species: the water hole.”

In 1992 NASA initiated a radio astronomy program called SETI.

In 1993 Congress cancelled the funding, claiming it was not valid science.



Mountainview, CA

The SETI Institute was incorporated as a 501(c)3 California Non-Profit Corporation on November 20, 1984. The inaugural officers of the Institute were CEO Thomas Pierson and SETI scientist Jill Tarter. The Institute began operations on February 1, 1985.

6 research divisions:

- Astrobiology
- Astronomy & Astrophysics
- Climate & Bioscience
- Exoplanets
- Planetary Exploration
- SETI



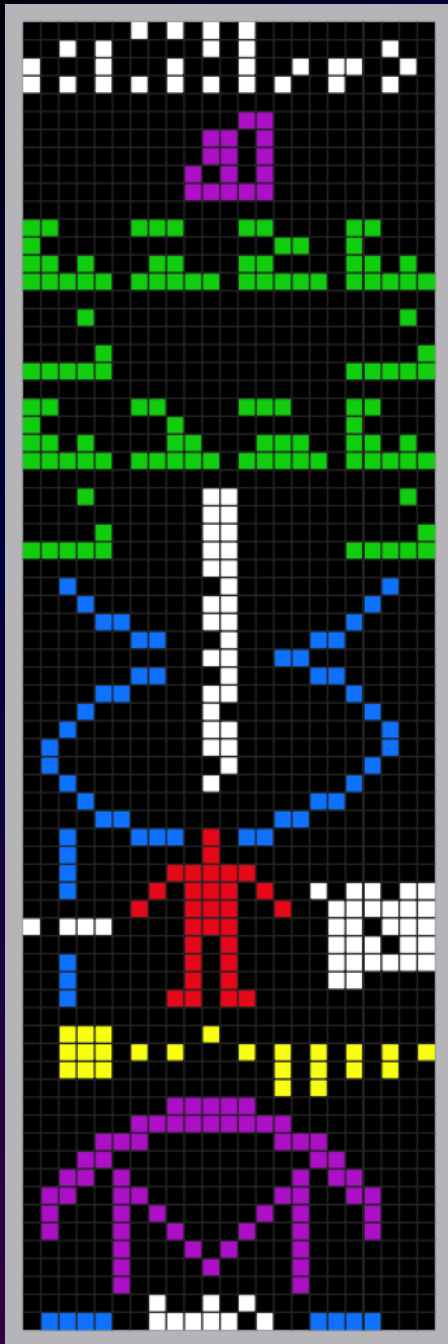
<https://seti.org/core-research>

and Frontier Development Lab

ARTIFICIAL INTELLIGENCE
RESEARCH FOR
SPACE EXPLORATION AND
ALL HUMANKIND



<https://frontierdevelopmentlab.org/>



What would the social, cultural, and political consequences be if we did make contact with an extra terrestrial intelligent civilization?

https://en.wikipedia.org/wiki/Potential_cultural_impact_of_extraterrestrial_contact